### COLLISIONLESS PHASE SPACE EQUILIBRIA IN PLASMAS

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### Goal

- Fusion Plasmas
- Solar Wind and most other astrophysical plasmas are weakly collisional (mean free path ~ 10<sup>8</sup> km)
- Measurements show uniformity and a Maxwellian-like velocity distribution
- When does the plasma equilibrate and how?
  - Simulate this in 1x1v and 2x2v

### **Statistical Mechanics**

- Phase Space (configuration space of position and velocity)
- Ergodic Hypothesis
  - Collisions
- Lynden-Bell Equilibrium



https://upload.wikimedia.org/wikipedia/commons/3/31/Ergodic\_hypothesis\_w\_reflecting\_rays.jpg

### Instabilities

- In 1d there are two-stream instabilities (generate electrostatic waves)
- In 2d there are Weibel instabilities (generate electromagnetic waves)





Initialized States for (a) 1d cold waterbag, (b) 1d hot waterbag, (c) 2d cold Maxwellian, (d) 2d hot Maxwellian

### Modelling Kinetic Plasmas with Gkeyll

- Discontinuous Galerkin
  - Finite element and finite volume
- Kinetic Modelling of collisionsless plasmas
  - Vlasov-Maxwell on a Cartesian grid
- Fitting simulation data with Stochastic Gradient Descent in PyTorch

$$\begin{split} \frac{\partial f_e}{\partial t} + \mathbf{v}_e \cdot \nabla f_e &- e\left(\mathbf{E} + \frac{\mathbf{v}_e}{c} \times \mathbf{B}\right) \cdot \frac{\partial f_e}{\partial \mathbf{p}} = 0\\ \frac{\partial f_i}{\partial t} + \mathbf{v}_i \cdot \nabla f_i + Z_i e\left(\mathbf{E} + \frac{\mathbf{v}_i}{c} \times \mathbf{B}\right) \cdot \frac{\partial f_i}{\partial \mathbf{p}} = 0\\ \nabla \times \mathbf{B} &= \frac{4\pi \mathbf{j}}{c} + \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t}\\ \nabla \times \mathbf{E} &= -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}\\ \nabla \cdot \mathbf{E} &= 4\pi \rho\\ \nabla \cdot \mathbf{B} &= 0 \end{split}$$

Vlasov-Maxwell Equations

### Collisions -> Exploring Phase Space -> Equilibrium Entropy is maximized

But in the Vlasov-Maxwell framework, there are no collisions and Entropy is a conserved quantity

# Goal is to Use Simulations to Investigate the of physics of collisionless equilibration



Probability Distribution Function plotted vs. position and velocity. Contours show the Hamiltonian isoenergic surfaces

Results (1x1v Phase Space)

Vlasov-Maxwell Simulation of the Two-Stream Instability using Gkeyll

### Fitting Distributions to the Data (1x1v)



 $logistic \ distribution = \frac{h}{1 - \exp(-k(x - 2(\mu - \sigma)))} + \frac{h}{1 - \exp(-k(x - 2(\mu + \sigma)))}$ 

### Fitting Distributions to the Data (2x2v)



#### **Beam Remnants**

- Fitted final few frames to a Maxwellian, and analyzed the residue from the fit, if it truly fits to a Maxwellian within an error, the residue should be a random distribution, but in fact, the residue is strongly correlated with the initial state, for the 1d simulation
- This contrasts with the 2d case where the residue from a Maxwellian fit is random

### **Conclusions and Relevance**

- Implies dimensionality plays a significant role in the relaxation of collisionless plasmas
  - Relaxation is more complete for 2d systems, may be due to the increased number of wave modes that are excited by 2d motion, electrostatic and electromagnetic wave modes
- Significant for theoretical work as they typically examine simplified 1d systems, such systems may not capture relaxation behavior of higher dimensional systems

### Acknowledgements

This work was made possible by funding from the Department of Energy for the Summer Undergraduate Laboratory Internship (SULI) program. This work is supported by the US DOE Contract No. DE-AC02-09CH11466

### THANK YOU!

Questions?

## SUPPLEMENTARY SLIDES

## Cold Water Simulation Video with Isoenergic Surfaces (1x1v)



## Hot Water Simulation Video with Isoenergic Surfaces (1x1v)



### Hot Water Simulation Video (2x2v)



### Computational Evidence Disagrees, There is Phase Space Mixing



#### Saturation and Equilibration (1x1v)



### **Two-Stream Instability**

![](_page_19_Figure_1.jpeg)

### About Me

- Rising Senior at the University of Pennsylvania majoring and submatriculating in Physics and minoring in Mathematics and German
- Working with Jason TenBarge at PPPL

### Results (1x1v)

![](_page_21_Figure_1.jpeg)